COL-OSSOS: Z-BAND PHOTOMETRY REVEALS THREE DISTINCT TNO SURFACE TYPES IDENTIFYING COLD CLASSICAL TNO SURFACES

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TNO SURFACES: WHAT & WHY THE KUIPER BELT Neptune

- TNOs were distributed into the outer Solar System during giant planet migration
- The orbital distribution of TNOs provides constraints on the dynamics of scattering
- The surface properties of TNOs provide constraints on the surface properties of the initial proto-planetesimal disk
- Surface properties result from composition and evolution



Minor Planet Center: Outer Solar System

TNO DYNAMICAL CLASSIFICATIONS



Gladman et al. 2008, SSBN

TNO SURFACES: WHAT & WHY



b. changing the planetary architecture



The Col-OSSOS Project

- Sample is from a characterized survey
- Follow-up photometry of a magnitude limited sub-sample of the Outer Solar System Origins Survey discoveries

Colossos Science Goals

- Determine the number of surface types
- Determine the intrinsic population fraction of these surface types
- Dependence between color and dynamics?
- Dependence between color and size?

Colors of the Outer Solar System Origins Survey (Col-OSSOS)





Col-OSSOS Photometry





z-Band Sample

Tegler & Romanishin 2003 Peixinho et al. 2003, 2012, 2015 Tegler et al. 2003, 2016 Barucci et al. 2005 Fraser & Brown 2012 Dalle Ore et al. 2013 Wong & Brown 2017

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z-Band Photometry

Cold Classical (CC)

Dynamically Excited

Blue Binary CC

5:1 Resonator

ightarrow

0

▲

Pike et al. 2017, AJ









Col-OSSOS Expansion Constraining Neptune's Migration: Surfaces of Resonant TNOs

Rosemary Pike, Kat Volk, Ruth Murray-Clay, Wes Fraser, Michele Bannister, Michael Marsset, Meg Schwamb, Matt Lehner, Aurelie Gulibert

- Dynamical simulations of Neptune's migration to predict the number of cold classicals in the 2:1
- Photometry of a complete sample of 2:1 TNOs from the OSSOS Ensemble characterized surveys in g, r, and z bands

K.Volk

Col-OSSOS Expansion Program on LBT

- LBT in Mount Graham, AZ
- 34TNOs in the 2:1, mr≤24.5
- z band using LBC Red
- g and r using LBC Blue

- 34TNOs in the 2:1, mr≤24.5
- April 2018: 2/3 nights lost to weather
- 5 2:1s and 2 Cold classicals observed
- 2018B: 3.5 nights awarded by NOAO

CONCLUSIONS

- Red Cold Classical TNO surfaces are distinct from dynamically excited TNO surfaces in the combination of g, r, and z-bands.
- Separating TNOs based on a bimodal distribution is not sufficient to correctly identify red and neutral TNO surfaces classes.
- Simultaneous or near-simultaneous photometry is necessary to distinguish the red dynamically excited TNOs from the neutral dynamically excited TNOs.
- Photometry in *grz* can be used to identify TNOs with cold classical surfaces outside the cold classical dynamical region, such as the 2:1 resonance. We target this resonance in our current LBT program.

COL-OSSOS RESULTS

Additional papers by the Col-OSSOS Team

TRIPPy: TRAILED IMAGE PHOTOMETRY IN PYTHON

Wesley Fraser¹, Mike Alexandersen², Megan E. Schwamb², Michaël Marsset³, Rosemary E. Pike⁴, J. J. Kavelaars⁵, Michele T. Bannister⁴, Susan Benecchi⁶, and Audrey Delsanti⁷

Figure 2. Example of a pill aperture. The full aperture, outlined in solid white, is the combination of a rectangle of length l, and width 2r, with two semicircular end-caps of radius r, all rotated at angle α . For aperture photometry, the background is measured inside a user specified box, and outside a pill aperture with larger r (but the same l). The image is a 480 s exposure of asteroid 2006, Polonskaya, taken on 2008 January 15 14:40:25 UTC when the asteroid had a rate of motion of $19^{\prime\prime}_{...}04 \text{ hr}^{-1}$ at angle 30°.9. The peak pixels are nearly 9000 ADU brighter than the background.

Star PSF subtraction

Asteroid PSF subtraction

Published in AJ 2016

All planetesimals born near the Kuiper belt formed as binaries

Wesley C. Fraser ➡, Michele T. Bannister, Rosemary E. Pike, Michael Marsset, Megan E. Schwamb, J. J. Kavelaars, Pedro Lacerda, David Nesvorný, Kathryn Volk, Audrey Delsanti, Susan Benecchi, Matthew J. Lehner, Keith Noll, Brett Gladman, Jean-Marc Petit, Stephen Gwyn, Ying-Tung Chen, Shiang-Yu Wang, Mike Alexandersen, Todd Burdullis, Scott Sheppard & Chad Trujillo

Published in Nature Astronomy, 2017

Gemini DD Award Calibrated images released to the public within 5 weeks of observations http://doi.org/10.11570/17.0010

First arrival from afar: the interstellar planetesimal 'Oumuamua

Michele Bannister, M. E. Schwamb, W. C. Fraser, M. Marsset, A. Fitzsimmons, S. Benecchi, P. Lacerda, R. E. Pike, JJ Kavelaars, A. B. Smith, S. O. Stewart, S.-Y. Wang (王祥宇), M. J. Lehner

Published in ApJL 2017

Col-OSSOS: Color and Inclination are Correlated Throughout the Kuiper Belt

Michael Marsset, Wesley C. Fraser, Rosemary E. Pike, Michele T. Bannister, Megan E. Schwamb, Kathryn Volk, J. J. Kavelaars, and OSSOS Core (In prep.)

"Blue" TNOs have larger orbital inclinations than "red" objects (dynamically hot populations) 50 ≥ 6 Blue class Red class - 5 40 -4 Inclination [°] 30 - 3 -2 10 - 1 125 62 $0 \downarrow -20$ 0 40 60 20 0 Spt. slope $[\%/(10^3 \text{\AA})]$

Implications: TNO surfaces are primordial and formed in a heterogeneous protoplanetary disk

Col-OSSOS: The Colours of the Outer Solar System Origins Survey

M. Schwamb, M. Bannister, W. Fraser, M. Marsset, R. Pike, J.J. Kavelaars, S. Benecchi, M. Lehner, S.Y. Wang, A. Thirouin, A. Delsanti, N. Peixinho, M. Alexandersen, Y.T. Chen, B. Gladman, S. Gwyn, J.M. Petit, K. Volk

Release of Photometry from:- H block (17 objects)- L block (18 objects)

Col-OSSOS Photometry

www.colossos.net

~THANKYOU~

Col-OSSOS: <u>www.colossos.net</u>

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