

RECENT GEMINI RESULTS ON X-RAY BINARIES

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GEMINI FOR XRBS

- Key questions:
Nature of XRB systems (e.g. ULXs)
Populations of XRBs
Masses of accreting objects
-Mass distributions illuminate SN physics
- Photometry, spectroscopy of donor stars
Use GMOS (imaging, spectra), GNIRS, Flamingos, NIFS

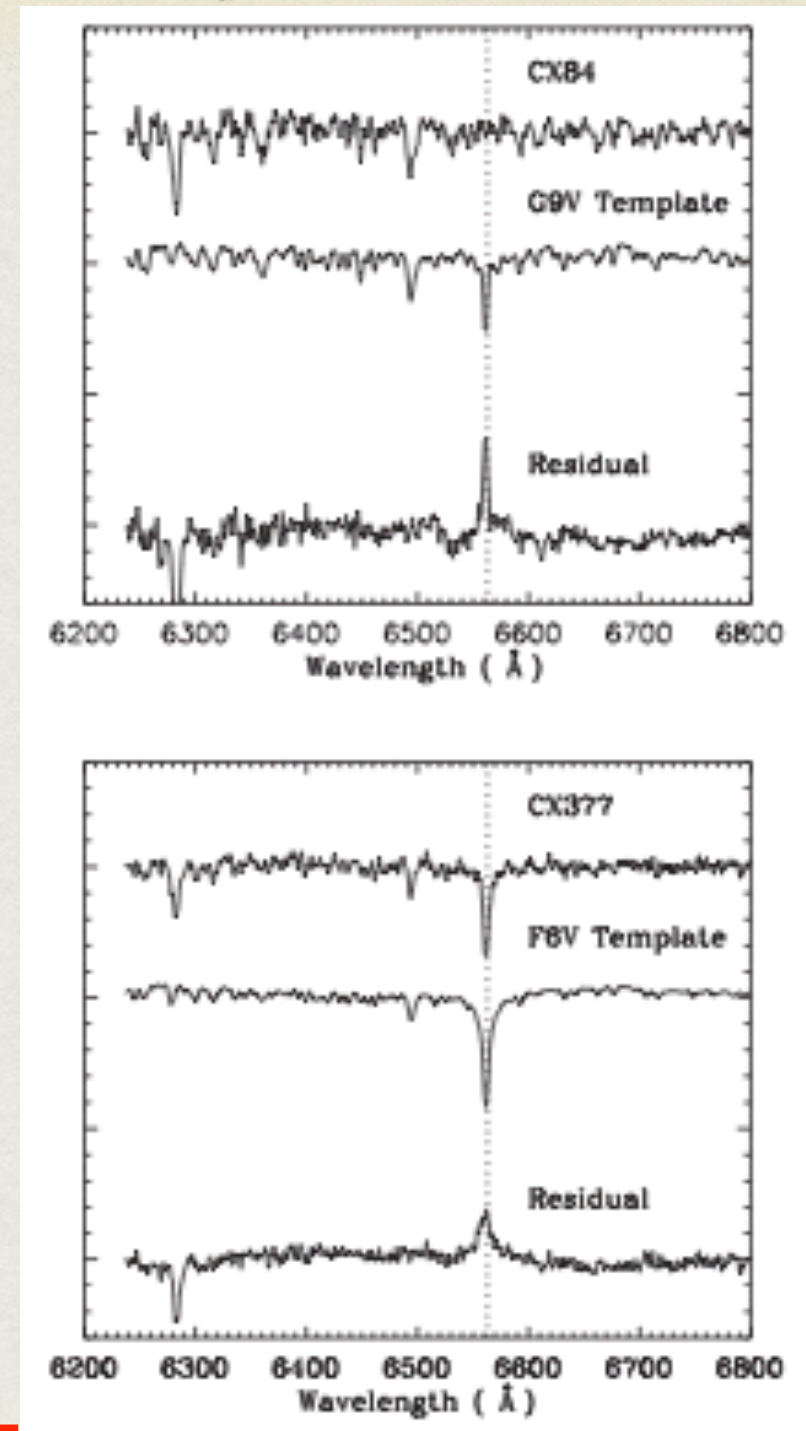


Courtesy ESA

HIDDEN ACCRETING BINARIES

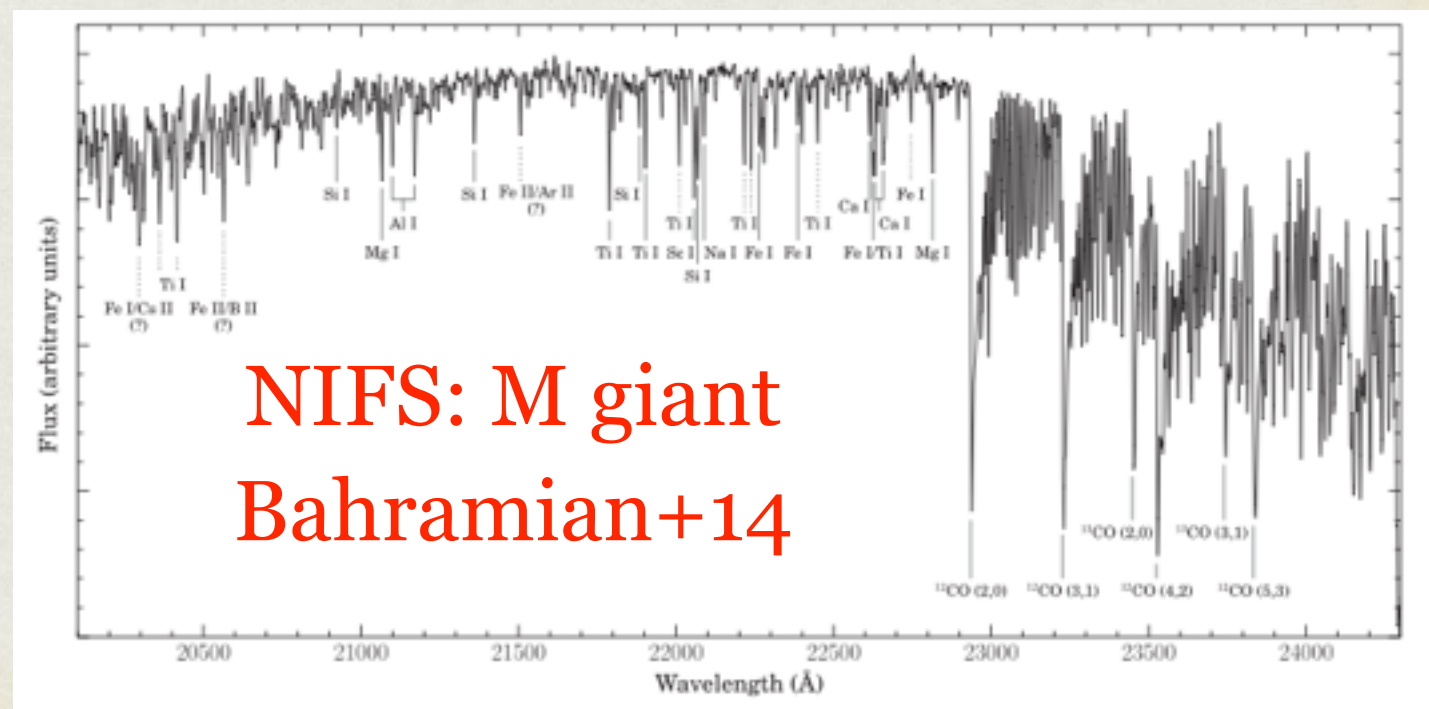
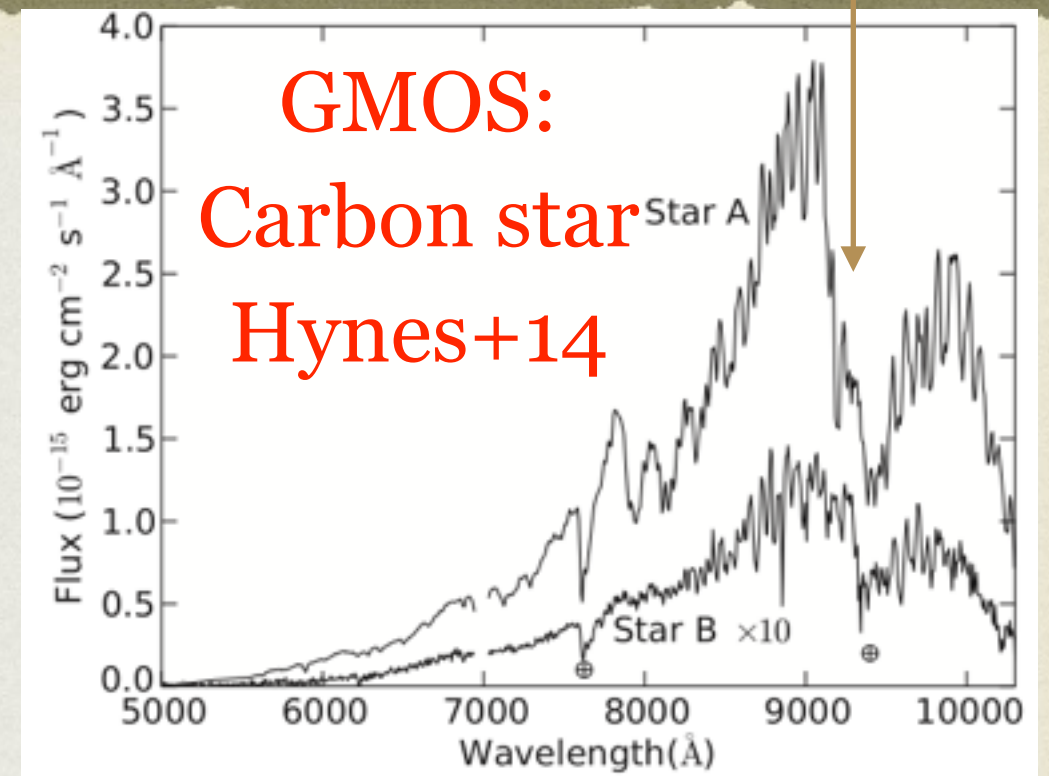
- Chandra Galactic Bulge Survey: 1640 sources (Jonker+11), follow-up opt/IR photometry, spectra to uncover ~ 300 XRBs
- GMOS data; some absorption-line spectra conceal broad emission lines (Wu+15)

GMOS spectra of
2 hidden XRBs, Wu+15



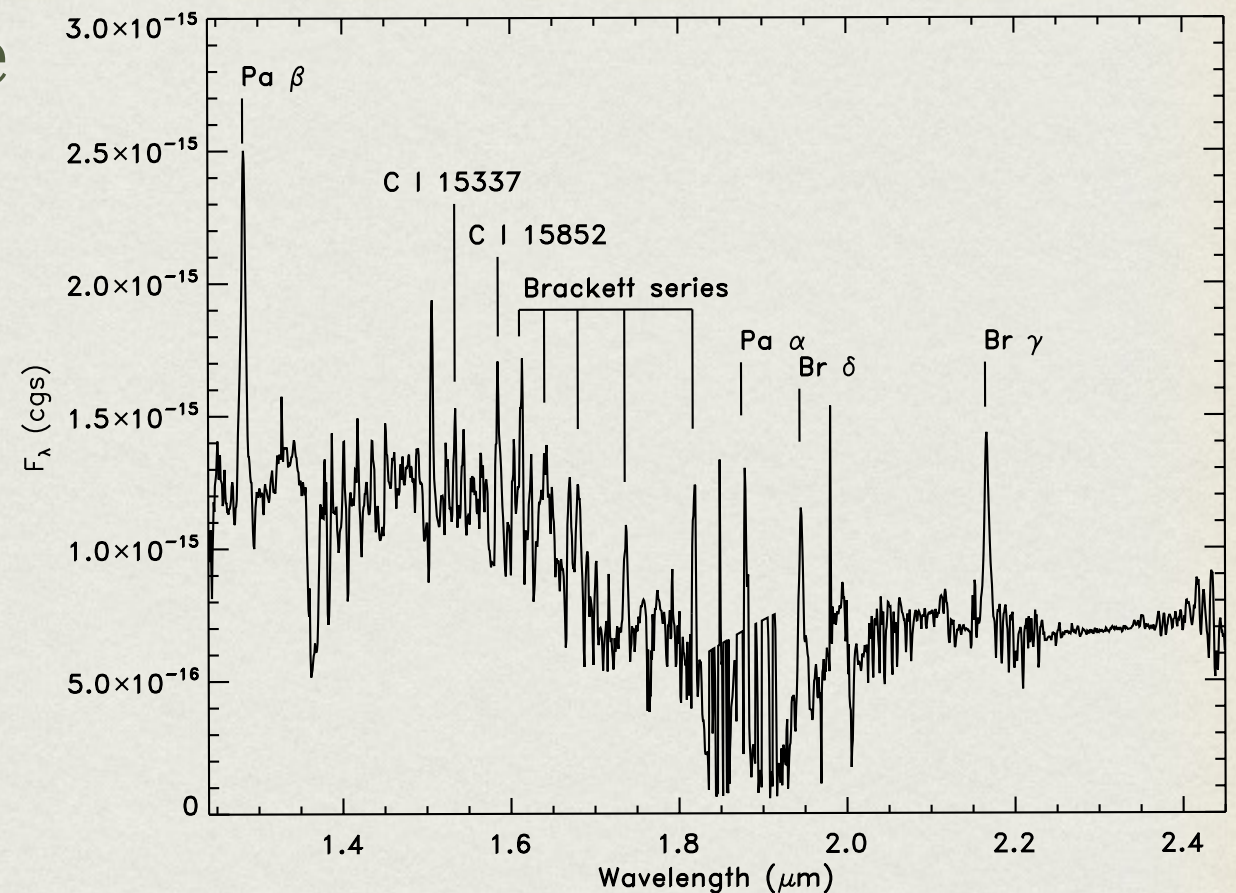
HIDDEN SYMBIOTIC X-RAY BINARIES

- Symbiotics are compact object accreting red giant wind
- 1000s predicted in galaxy, only ~100 known. Most don't show emission lines? (Hynes+14)
- Gemini to type giant, measure distance & infer accretor nature (e.g., L too high for white dwarf)



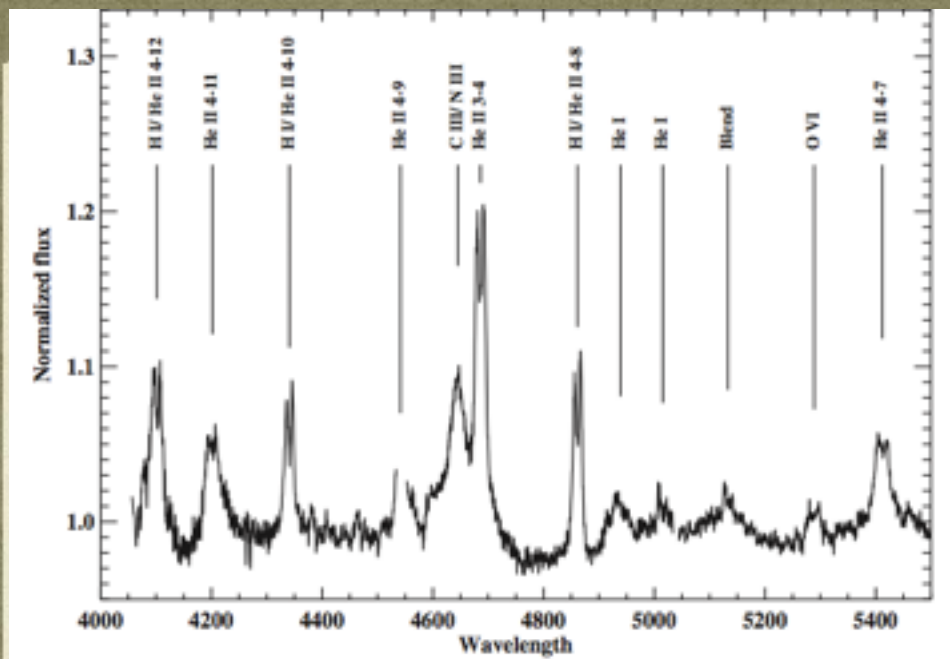
ACCRETING PROTOSTAR?

- Faint X-ray source from Bulge survey
Red SED (peak at $10\ \mu\text{m}$)
Mid-IR outburst >6 mag
Emission lines, including forbidden
- Interpret as young stellar object (FU Ori?).
3 degrees above plane;
isolated star formation?



Flamingos-2 on CX330,
Britt+15, submitted
(GRACES test target)

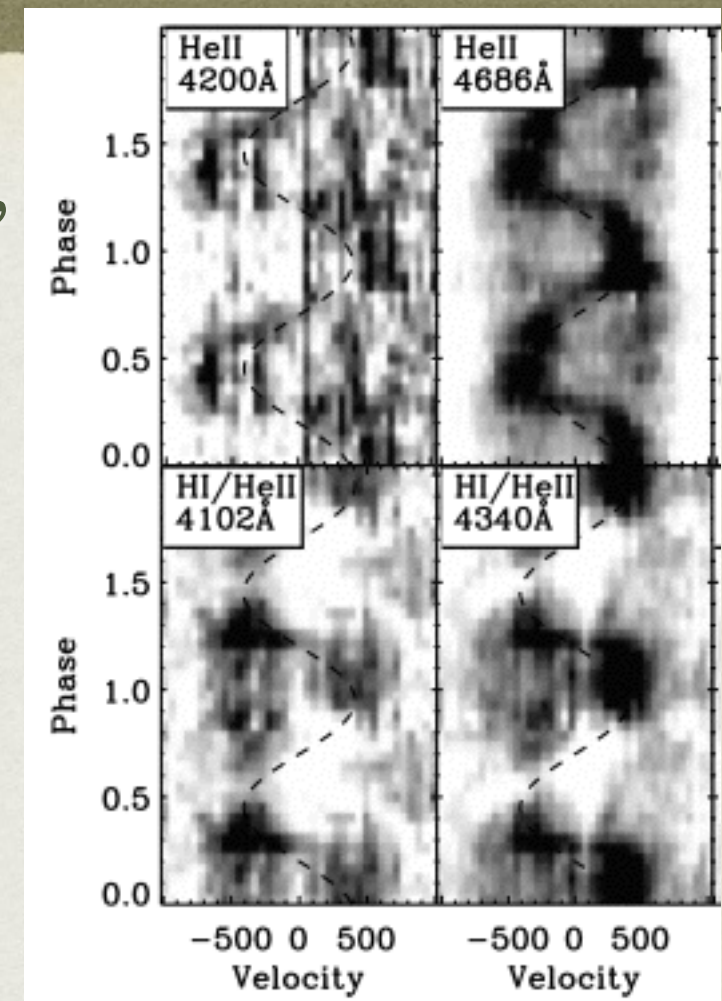
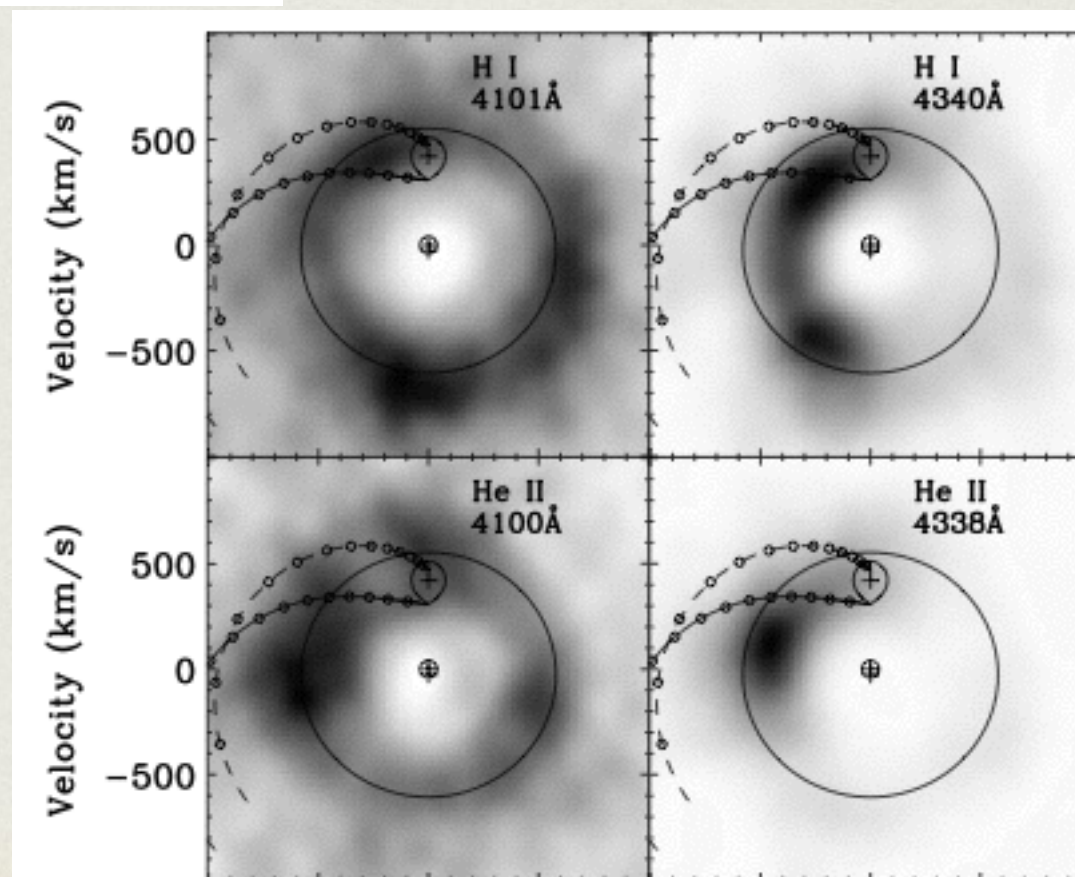
RADIAL VELOCITY STUDIES



Spectrum, with
double-peaked
emission lines

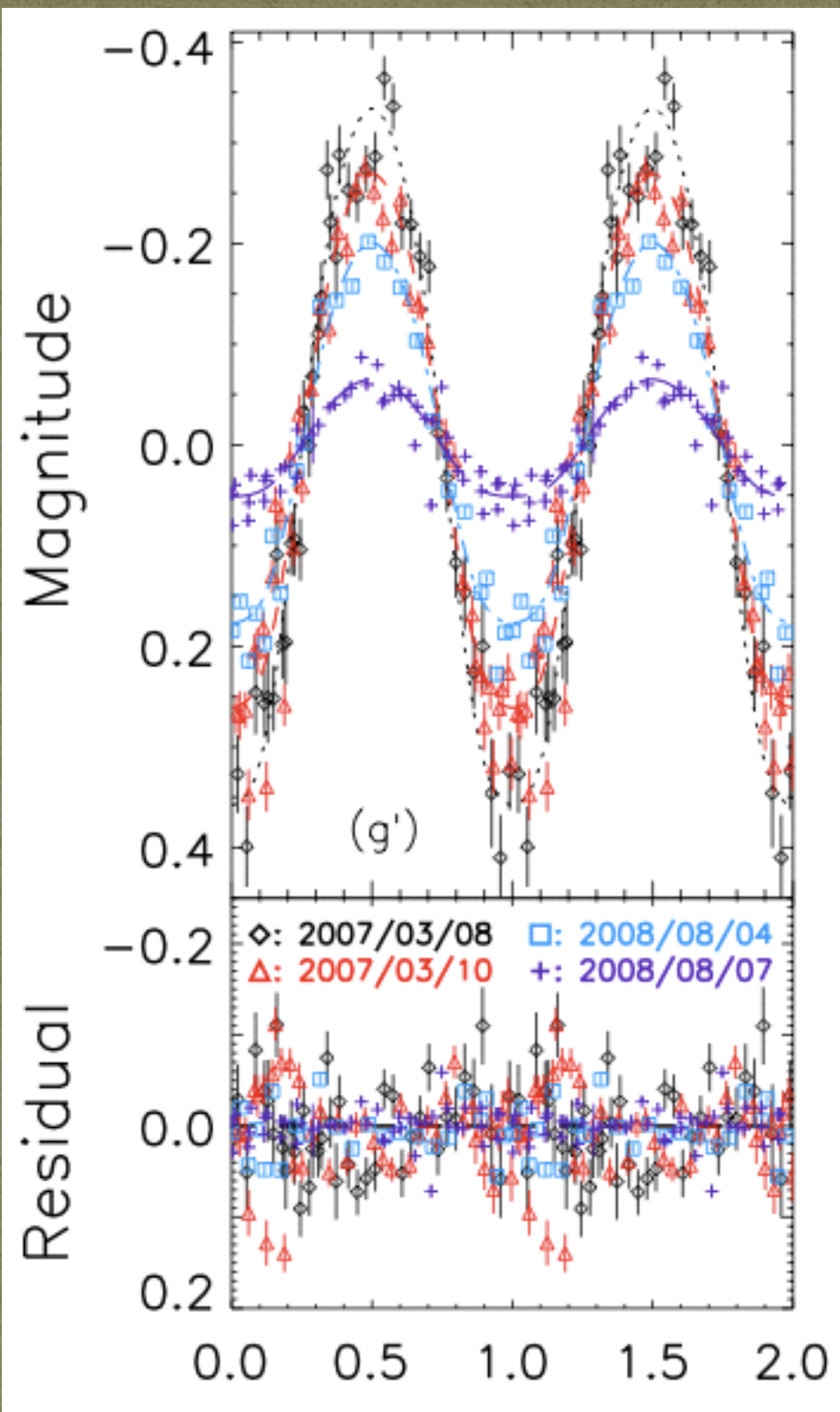
EXO 0748-676,
Mikles+12

Trailed spectra,
change with
orbital phase



Tomograms, show
locations of emission

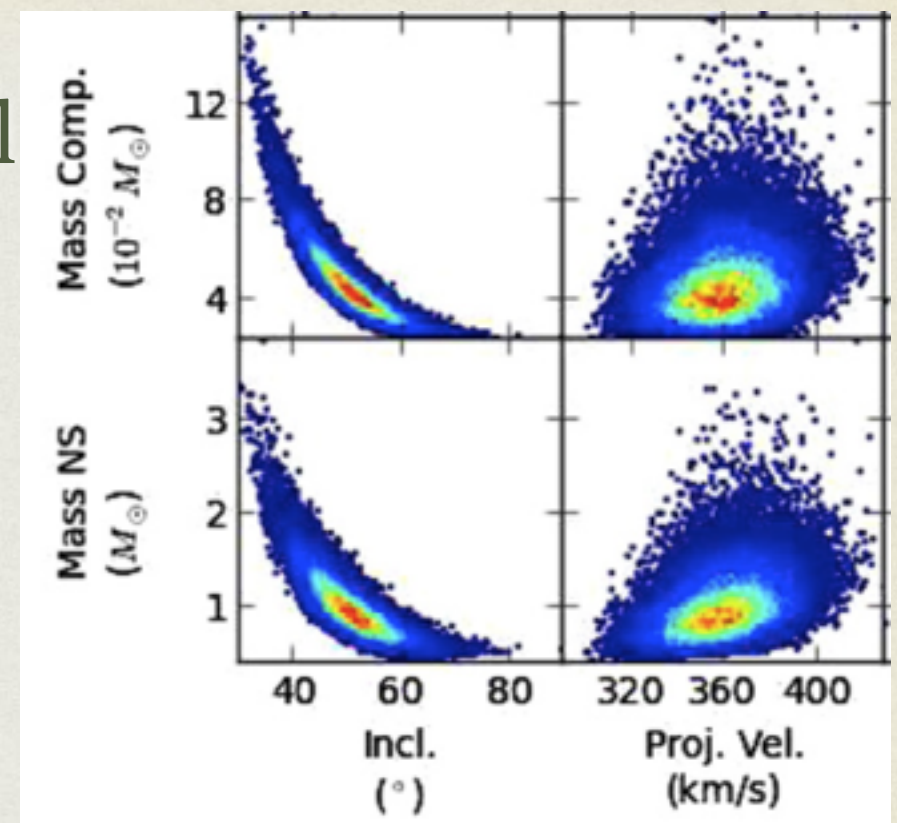
SPECTRA + PHOTOMETRY



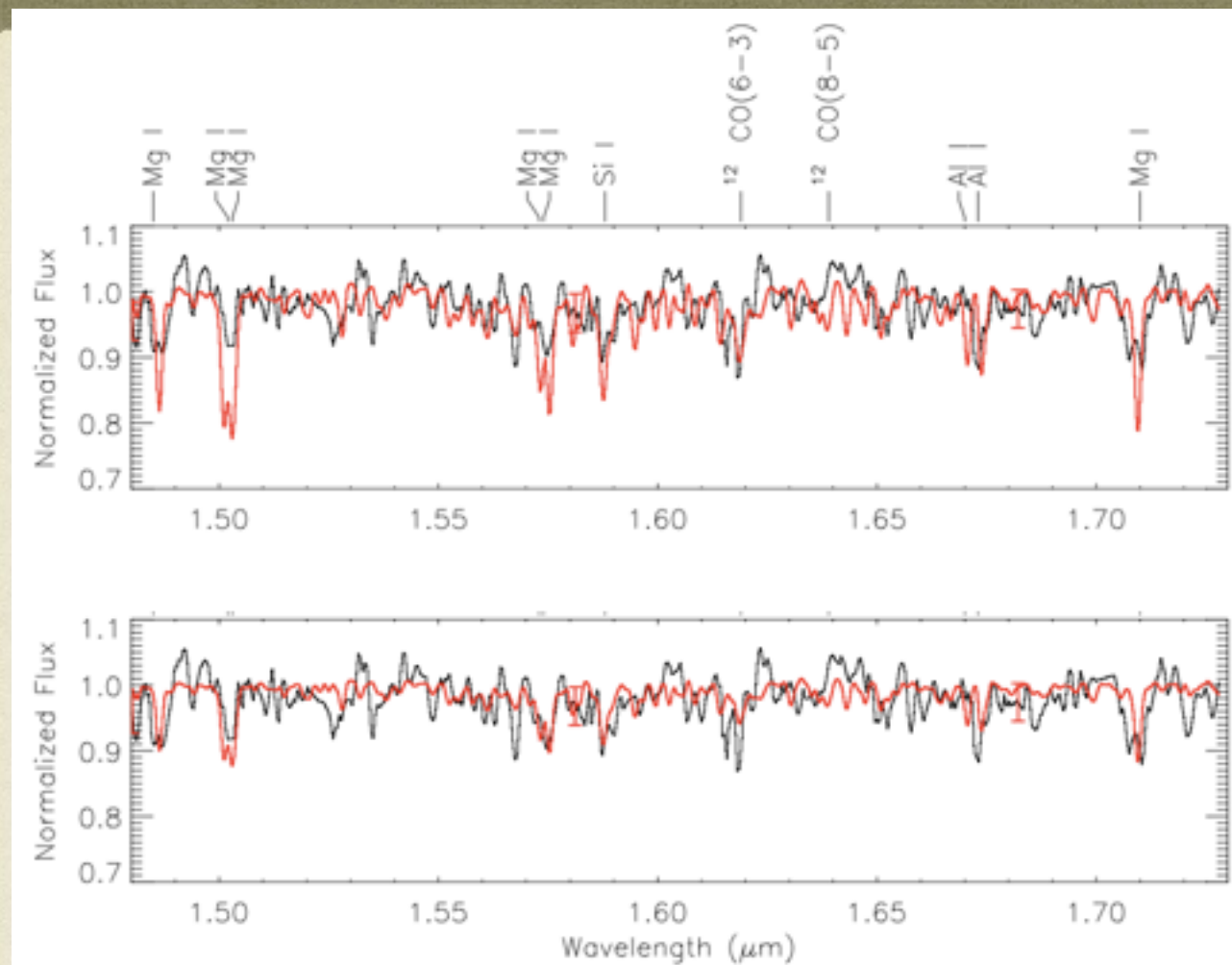
Degeneracy
between radial
velocities &
inclination

Multi-band
light curves
constrain
ellipsoidal
variations,
inclination

SAX J1808-37;
Wang+13



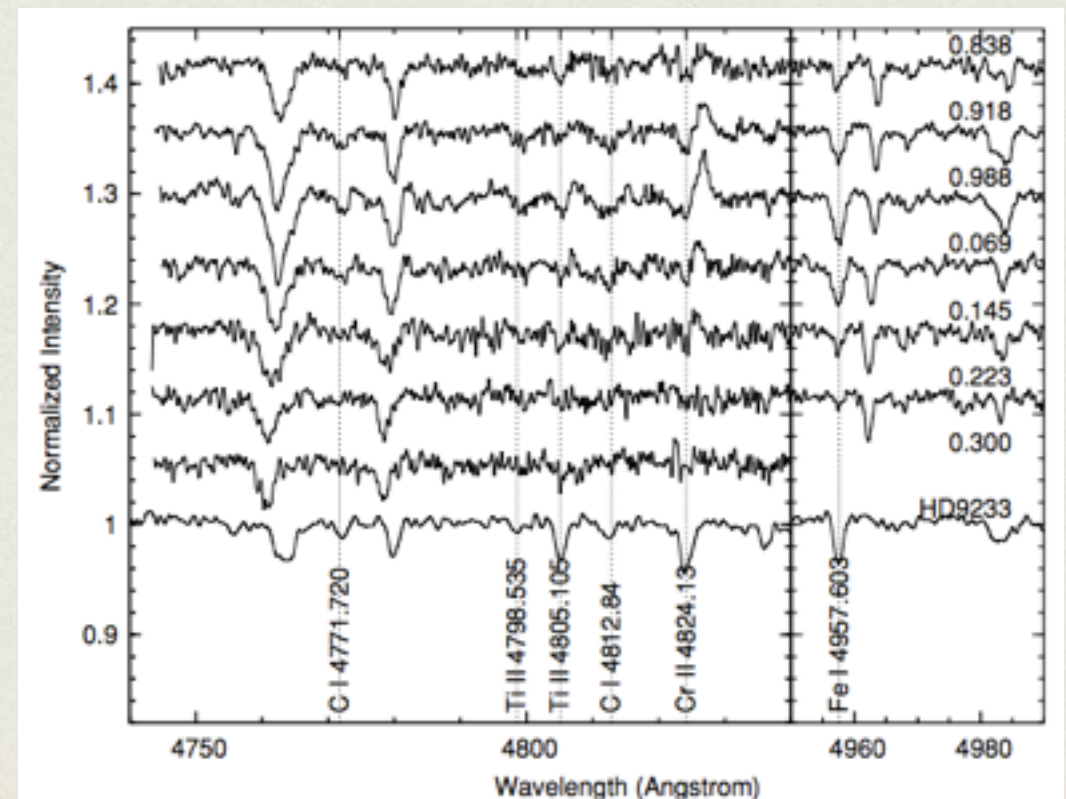
MEASURE BLACK HOLE MASSES



GNIRS H-band spectrum, half of IR light from disk in XTE J1118+480. With photometry, allows $6.9 < M_{\text{BH}} < 8.2$ Msun.

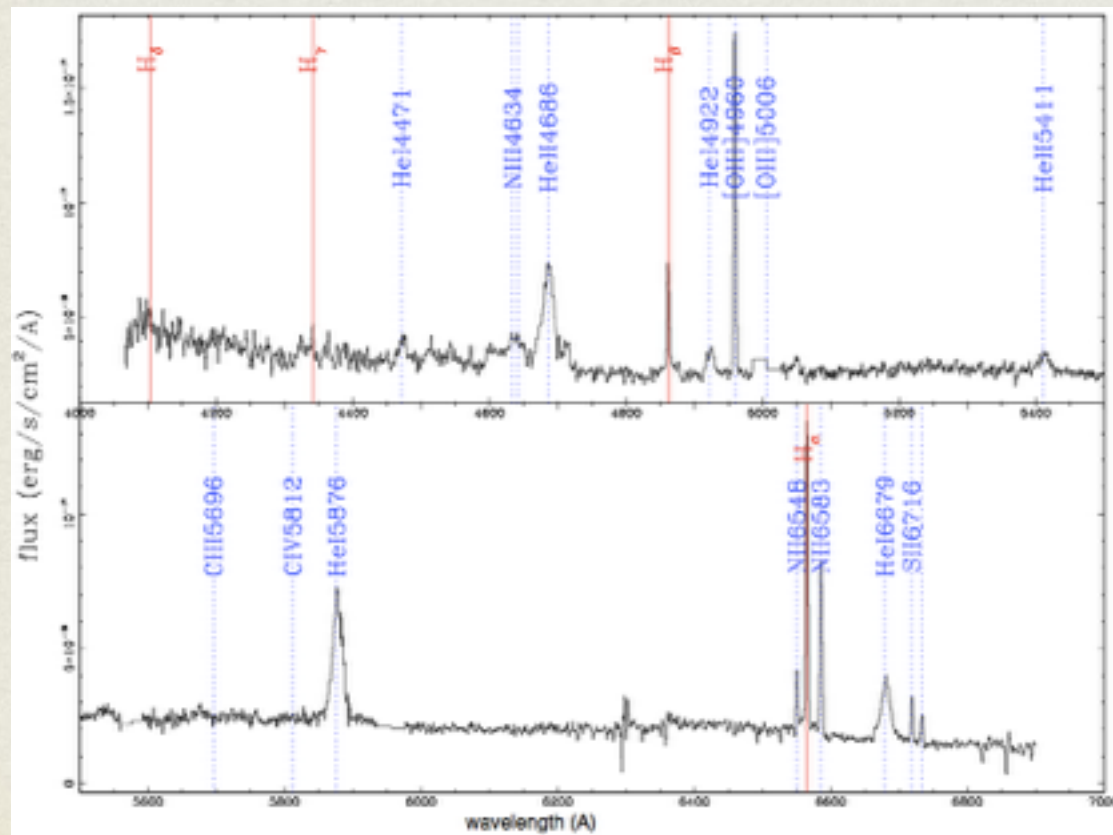
Khargharia+13

Complex GMOS spectra of SS433 give RVs, giving $1.9 < M_{\text{BH}} < 4.9$ Msun. Kubota+10

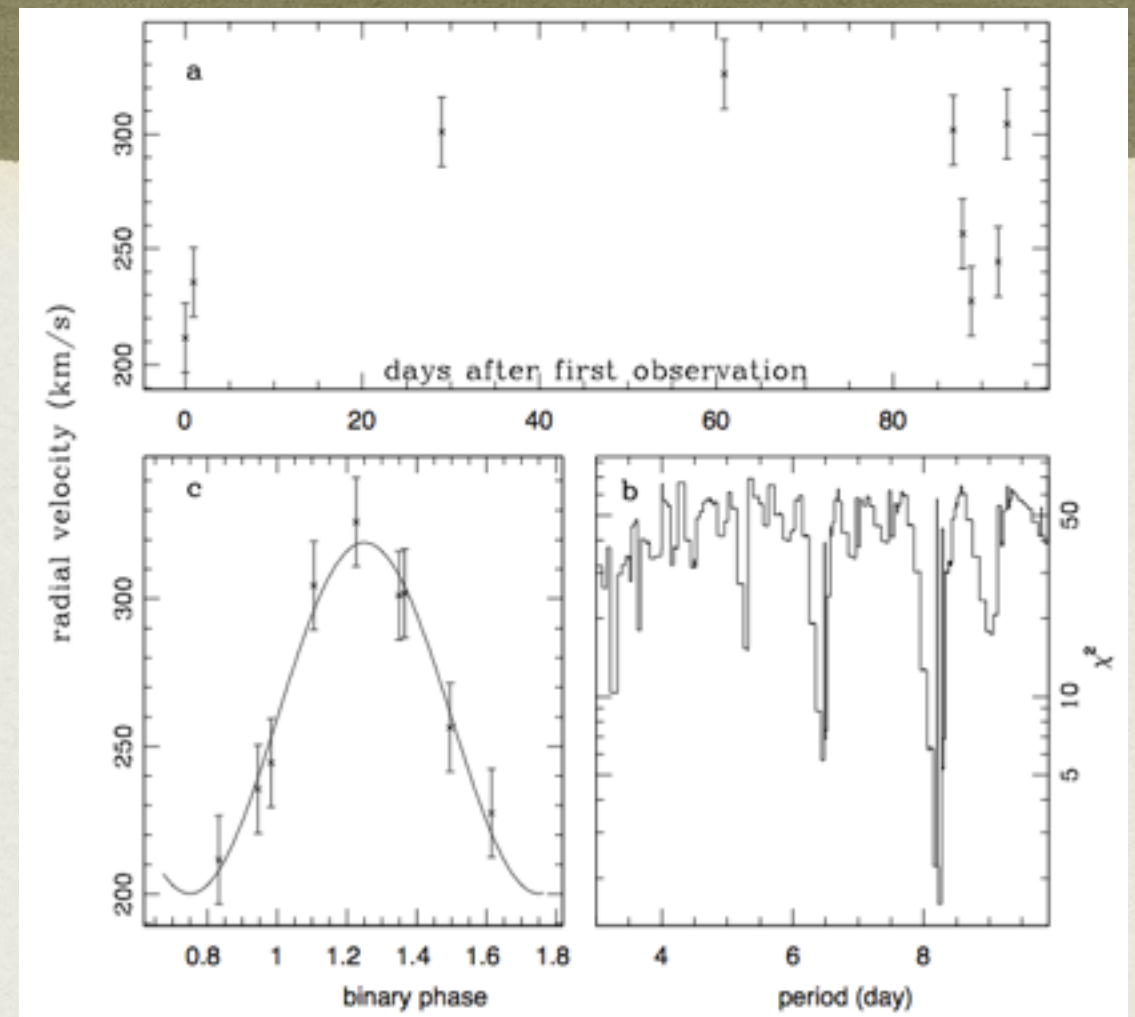


MASS OF BH IN A ULX

GMOS identified bright XRB (ULX) in M101 with WR star, gave mass estimate for star, 19 Msun (Liu+13).



Liu
+13



GMOS radial velocities gave orbit, BH mass 20-30 Msun.

Proves some of brightest XRBs are **not** ~1000-Msun BHs.

GALACTIC BULGE SURVEY LLP

- Approved LLP (PI R. Hynes);
GMOS dynamical mass estimates of ten quiescent XRBs in Galactic Bulge.
Known systems selected in outburst—selection effects?
Measure mass distributions of NSs & BHs,
constrain SN explosion mechanisms.